

NASA'S JOURNEY TO

# MARS

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**Michael Meyer, Lead Scientist**  
Mars Exploration Program, HQ

**MEPAG**  
**Feb. 2015**

# Mars Exploration

Operational

2001 - 2014

2016

2018

2020

2022



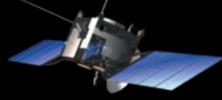
Odyssey



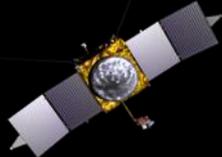
MRO



Mars Express  
Collaboration



ISRO - MOM  
Mangalyaan



MAVEN



ESA Trace Gas  
Orbiter  
(Electra)



Opportunity



Curiosity -  
Mars Science  
Laboratory



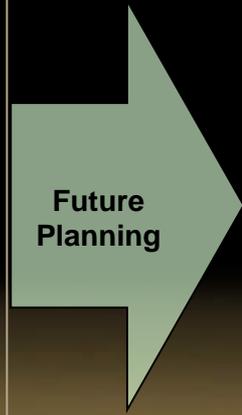
InSIGHT  
(Discovery Program)



ESA ExoMars  
Rover (MOMA)



2020  
Mars Rover



Future  
Planning

# Mars Exploration Program Analysis Group (MEPAG)

chartered by NASA HQ to assist in planning the scientific exploration of Mars



## Events

### ✓ **Status of Mars Missions**

- ODY, MER-B, MRO, MEX/NASA, MSL extended missions being implemented
- MAVEN and MOM in their science orbits around Mars and conducting their prime science missions (reports on Day 2)
- All Mars missions survived Comet Siding Spring encounter with Mars!

### ✓ **Science Highlights**

- Comet Siding Spring observations and impacts on Mars atmosphere
- *Curiosity* on slopes of Mt. Sharp, drilling away
- *Opportunity* examines rock diversity on Endeavour Crater rim; nearing marathon run.
- *Odyssey* getting a look at the morning/evening atmosphere at their new local time

### • **Progress being made on Landing Site Observations** (see talks on Day 2)

### • **Many science presentations at DPS, GSA, and AGU**

### • **Looking Ahead**

- ISRO-NASA Mars Working Group met face to face in Bangalore
- Face to Face MEPAG meeting on Feb. 24-25, 2015, Pasadena area.
- LPSC coming up on March 15-20
- FY17 Budget in work

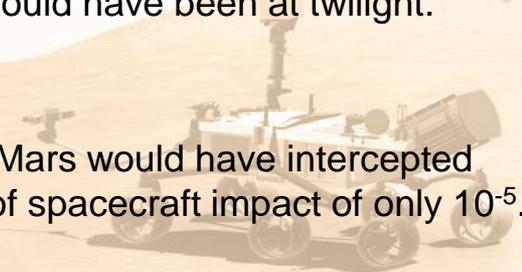




## Comet Siding Spring Highlights (observed at Mars)

- All spacecraft survived with no apparent damage!

- Comet:
  - Scale: ~ 1.6-cm wide brushed area
  - Scale: ~ 6-cm wide brushed area
  - **First-ever imaging of the nucleus of a long-period comet:**
    - MRO HiRISE: nucleus  $\leq 1$  km in diameter, may be  $\leq 500$  m (pixelated bright area)
    - Rotation and diameter determinations complicated by bright jets and “half-moon” phase
      - But nucleus is smaller than expected and it is rotating!
  - **Composition:**
    - MAVEN IUVS and NGIMS: First ever determination of metals in dust from an Oort Cloud comet.
- **Interaction with Mars:**
  - **Meteors did enter Mars atmosphere!**
    - Temporary detection of metal ions in substantial quantity
    - Temporary, but significant, ionospheric enhancement seen near ~100 km altitude (too deep for molecular penetration)
    - Rovers did not attempt to see meteor shower which would have been at twilight.
  - **How many meteors entered Mars atmosphere?**
    - Seems to have been a lot of material
    - Note: Even at an estimated fluence of 1 particle/km<sup>2</sup> Mars would have intercepted millions of meteors (> 1 metric ton) with a probability of spacecraft impact of only  $10^{-5}$ .

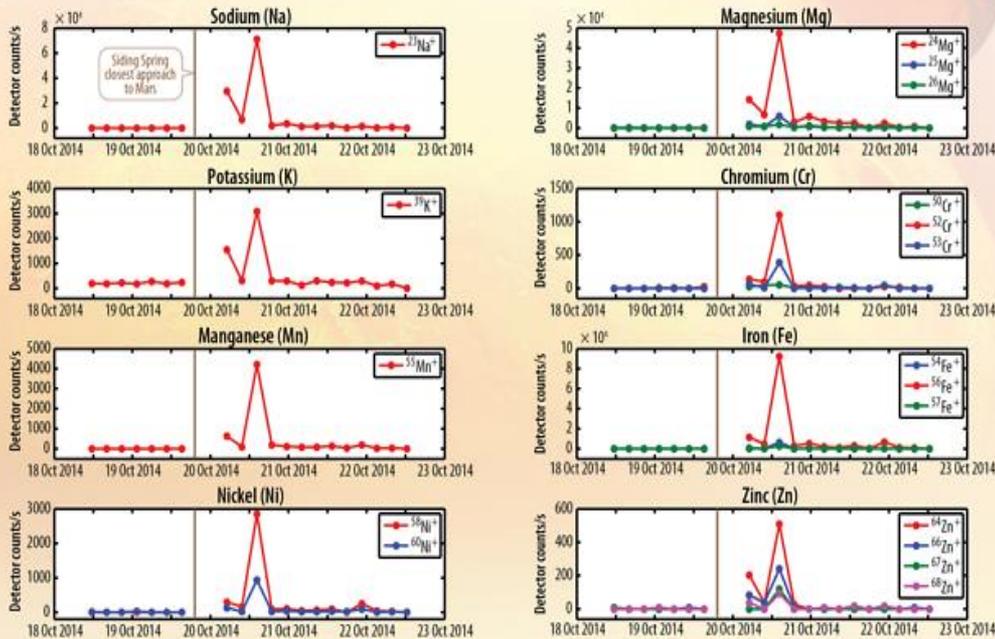


# Comet /2013 A1 Siding Spring: Comet Dust in the Mars Atmosphere (see MAVEN presentation)

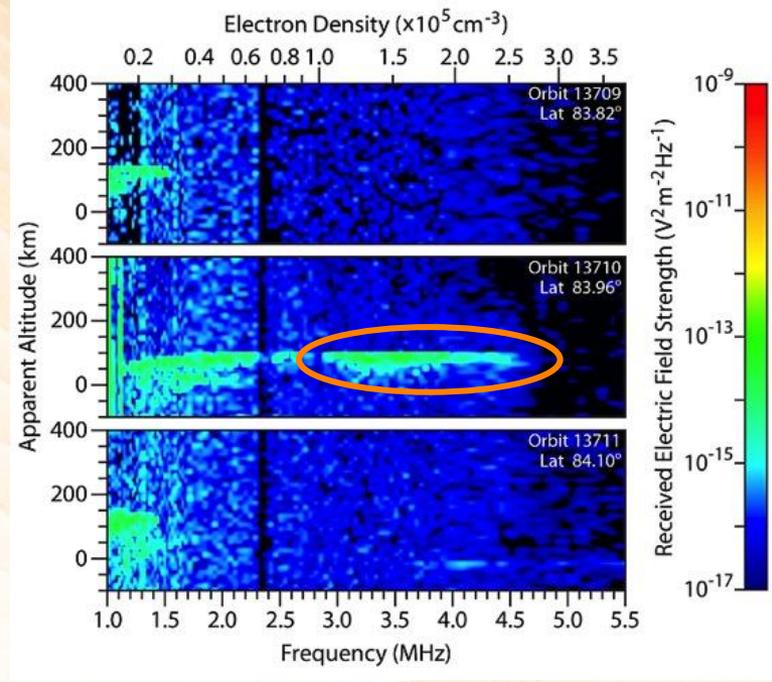
**MAVEN IUVS & NGIMS:**  
Metallic signatures of Comet Dust in the Mars Atmosphere

**MEX MARSIS & MRO SHARAD:**  
Comet-induced Ionization in the Mars Atmosphere

Eight different metal ions from comet Siding Spring were detected by NGIMS



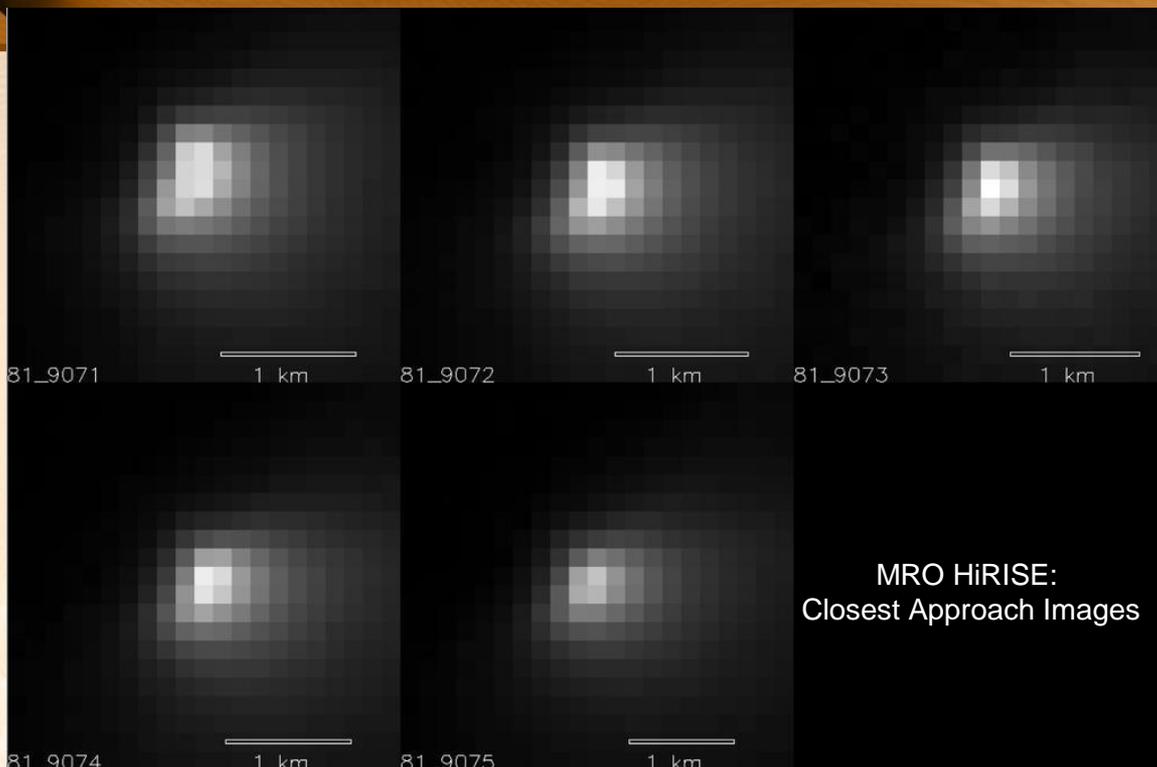
**MAVEN NGIMS:** In situ metal detection (~175 km)



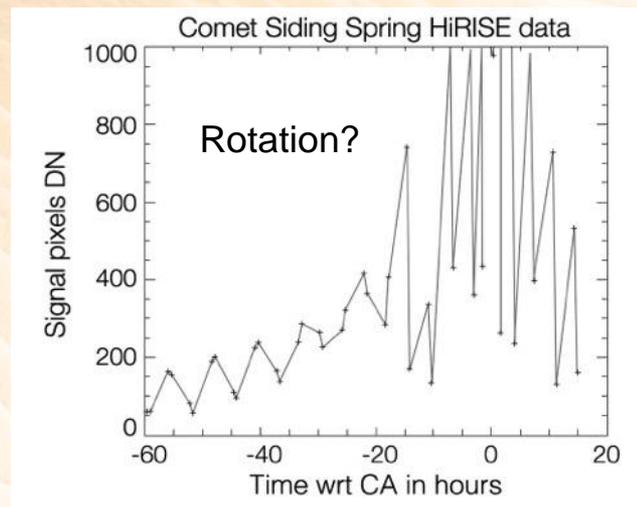
**MEX MARSIS:** Transient Ionospheric Layer at ~100 km



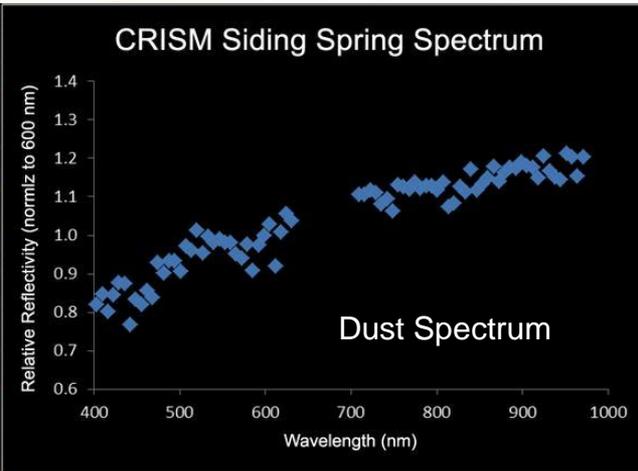
# Comet /2013 A1 Siding Spring: Nucleus & Inner Coma



**MRO HiRISE:**  
Brightness variation during CSS Encounter



**MRO CRISM:**  
*Left:* Comet Siding Spring Color Composite From Closest Approach (CA) Images  
*Right:* Composite Spectrum of CSS Reflecting Dust



## Science Highlight

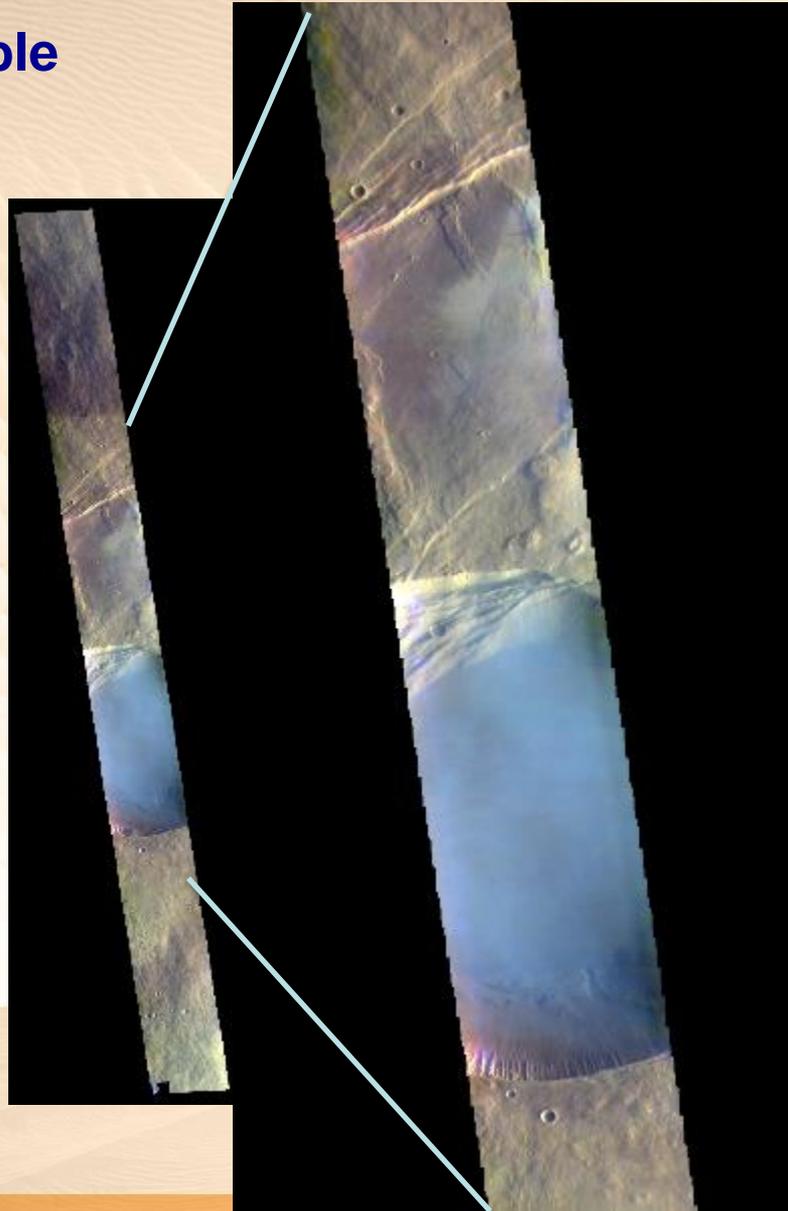
### **THEMIS Morning Day-Side Visible Images of Water-Ice Clouds**

Recent THEMIS data provide a preview of the morning day-side orbit geometry that Odyssey will be in for much of the upcoming Extended Mission 6.

This enhanced visible color image shows water-ice clouds that formed overnight lingering in the caldera of Pavonis Mons, near the Martian equator.

Image width 18 km.

Image acquired Sep. 4, 2014.



# Mars Exploration Program Analysis Group

chartered by NASA HQ to assist in planning the scientific exploration of Mars



## Landing Sites

### MRO

*Scale: ~ 1.6-cm wide brushed area*

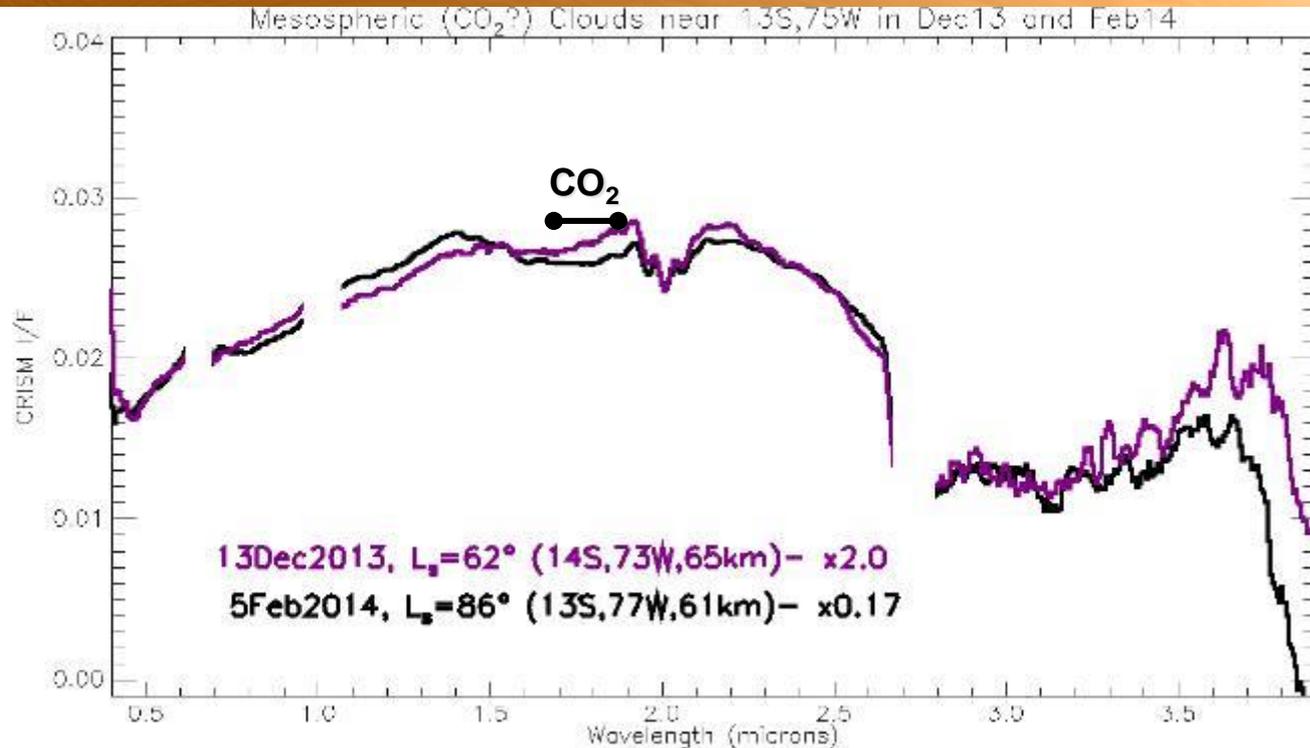
*Scale: ~ 6-cm wide brushed area*

- MRO has satisfied >50% of all current landing site characterization requests for InSight, 2018 ExoMars, & 2020 Mars rover
- Pace of MRO data acquisition for landing site characterization was slowed temporarily due to Comet Siding Spring activities and November “no roll” period. MRO now back to observing with nominal  $\pm 30^\circ$  rolls
- **Site Selection Activities: InSight, ExoMars, 2020 Mars Rover, Humans to Mars**
  - See specific presentations, many on Day 2.
  - 2020 *Second* Landing Site Workshop scheduled for first week in August 2015
  - Site proposers were asked to identify “regions of interest” for their site to aid 2020 Project assessment of operability.
    - May also benefit studies of landing site configuration for humans



# High Altitude CO<sub>2</sub> Ice Clouds on Mars

61 km altitude  
13°S, 77°W

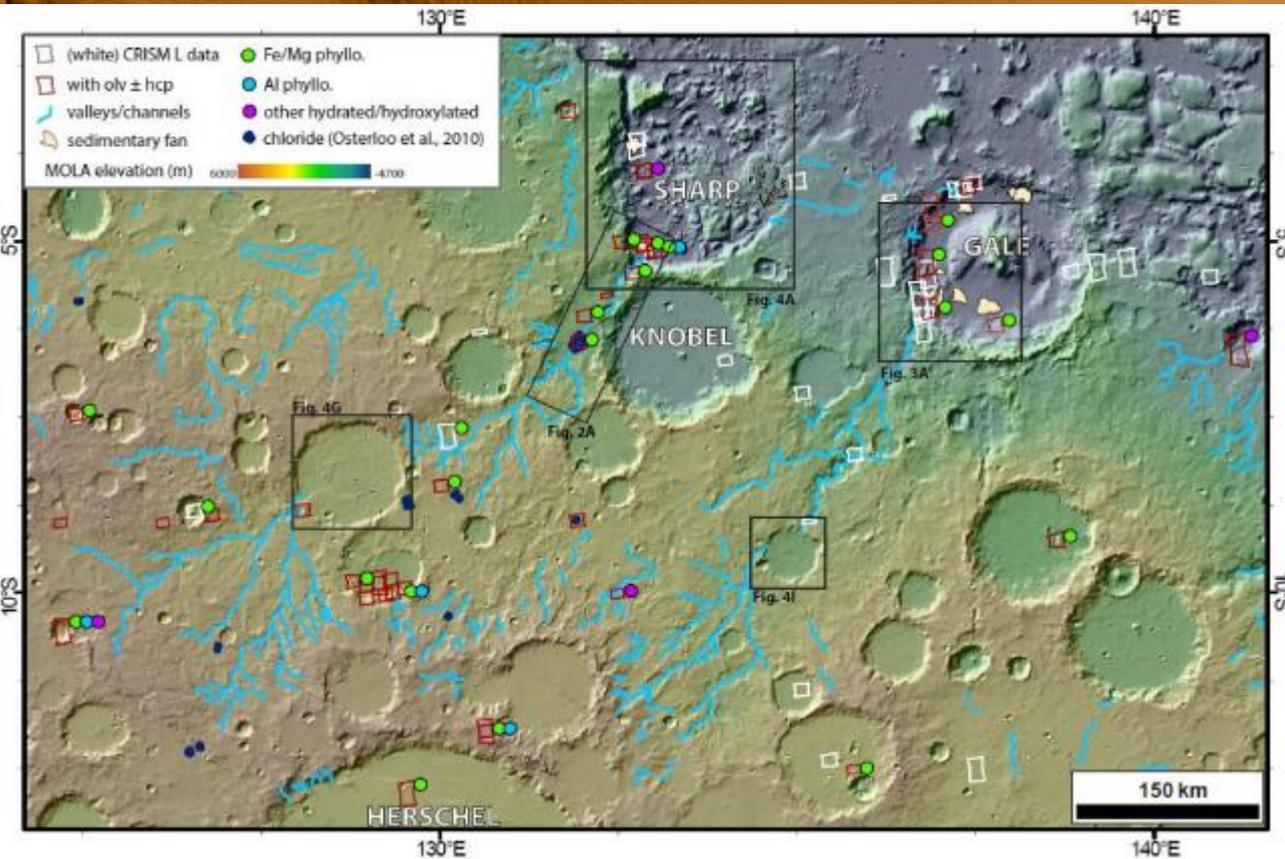


**Left:** False color limb scan  
(R= 2.53, G=1.51, B=1.08 μm)

**Above:** stretched using CO<sub>2</sub>  
bands (R=2.14, G=2.00,  
B=1.87 μm)

- **Mission/instrument:** MRO/CRISM
- **Finding:** Rare, high altitude CO<sub>2</sub> ice clouds occur in Mars' northern equatorial region of during northern summer/fall.
- **Importance:** CO<sub>2</sub> ice clouds like these are thought to have had a strong influence on climate early in Mars' history.
- **Reference:** Clancy, R. T., et al. (2014). CRISM Limb Observations of Mars Mesospheric Ice Clouds: Two New Results. Lunar and Plan. Sci. Conf. Abstract #1006.

# Minerals in Watersheds Around Gale Crater



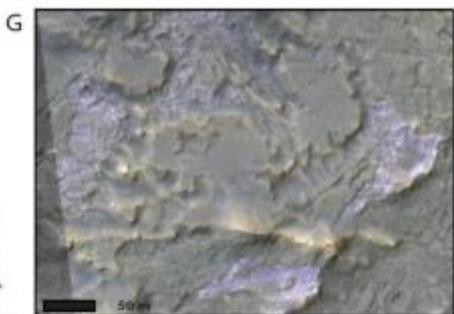
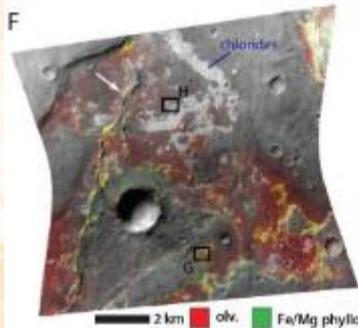
■ **Mission/instrument:**  
MRO/CRISM

■ **Finding:** Gale and surrounding craters host similar bedrock mineralogy and altered clay deposits on their floors, but vary in evaporate salt content.

■ **Importance:** Fluvial processes emplaced detrital clay minerals. As Mars dried out, the three craters became hydrologically isolated and salts were left behind nonuniformly.

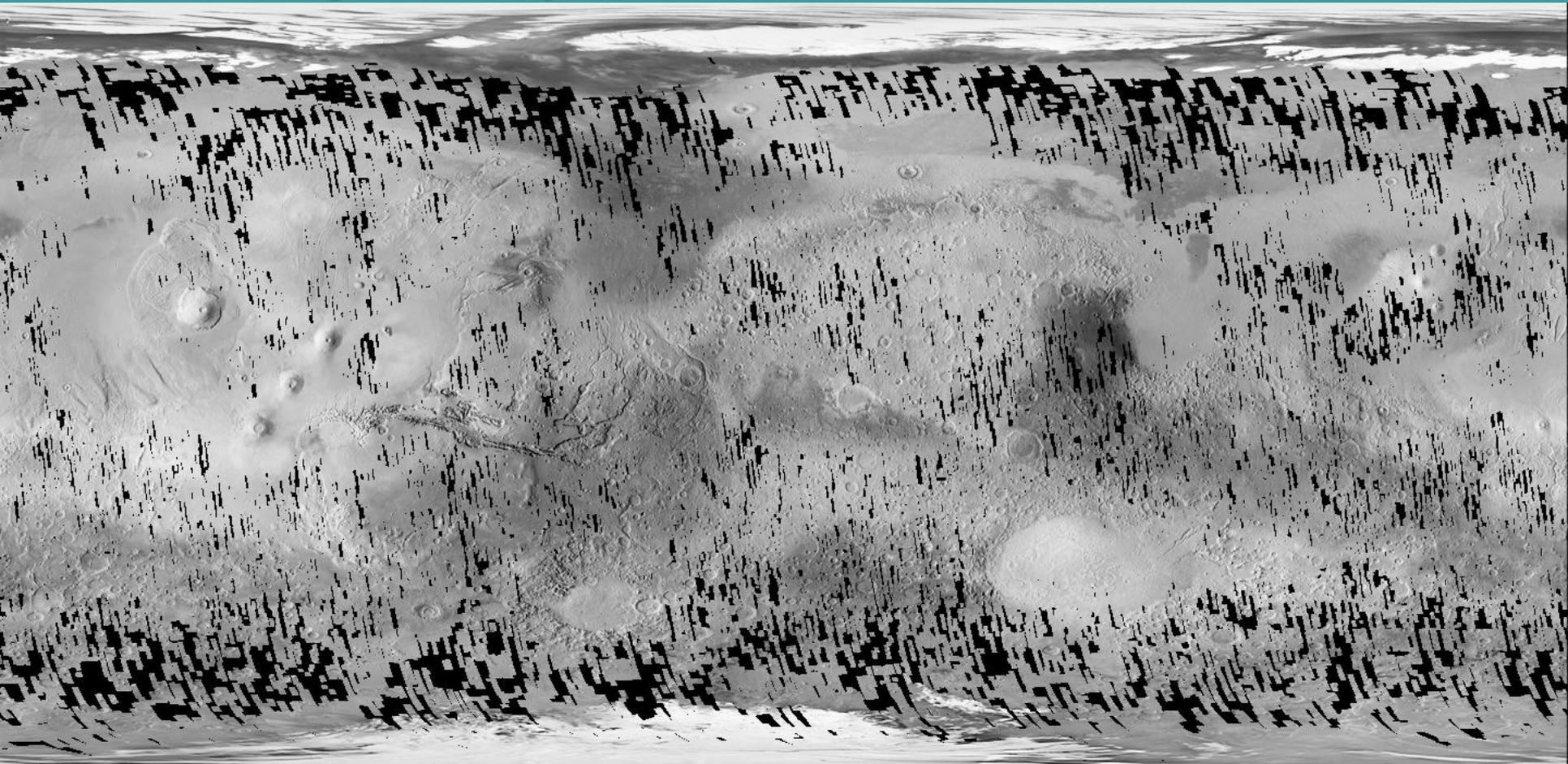
Curiosity will explore this

**geologic history.**  
 Reference: Ehlken, B. L., and Buz, J. (2014) Mineralogy and Fluvial History of the Watersheds of Gale, Knobel, and Sharp craters: Context for MSL Curiosity's Exploration, Geophys. Res. Lett., DOI: 10.1002/2014GL062553.



# CTX Coverage

△ Coverage: January 2013 – 2015



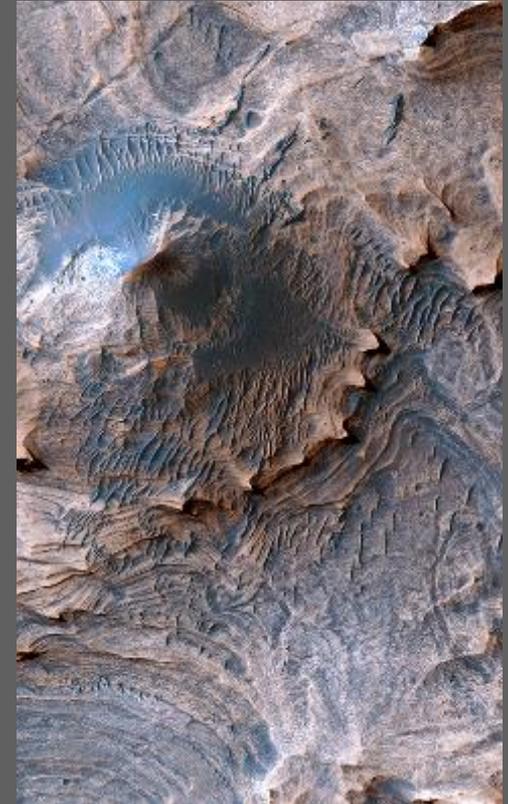
CTX has now covered ~94% of Mars

# Marsquakes and water-lain sediments in Candor Chasma



Perspective view of the Candor Colles region of west Candor Chasma, Mars. HiRISE image PSP\_001641\_1735

- Unprecedented high-resolution geologic and structural mapping using HiRISE data has revealed evidence of past marsquakes and water-lain sediments in the west Candor Chasma region of Valles Marineris.
- Orientations of the sediment layers indicate that these rocks formed as sand and dust was blown in by the wind and became trapped in shallow playa lakes.
- Injectite megapipes are also observed. These features formed by underground movement of the water-lain sediments in response to ground shaking (marsquakes) from several large fault zones in the area.
- These injectites served as reservoirs for groundwater and thereby would have hosted potentially habitable environments in the Martian subsurface approximately 3 to 3.5 billion years ago.
- Results published in Okubo, C.H., 2014, Bedrock geologic and structural map through the western Candor Colles region of Mars: U.S. Geological Survey SIM 3309, scale 1:18,000, <http://dx.doi.org/10.3133/sim3309>.

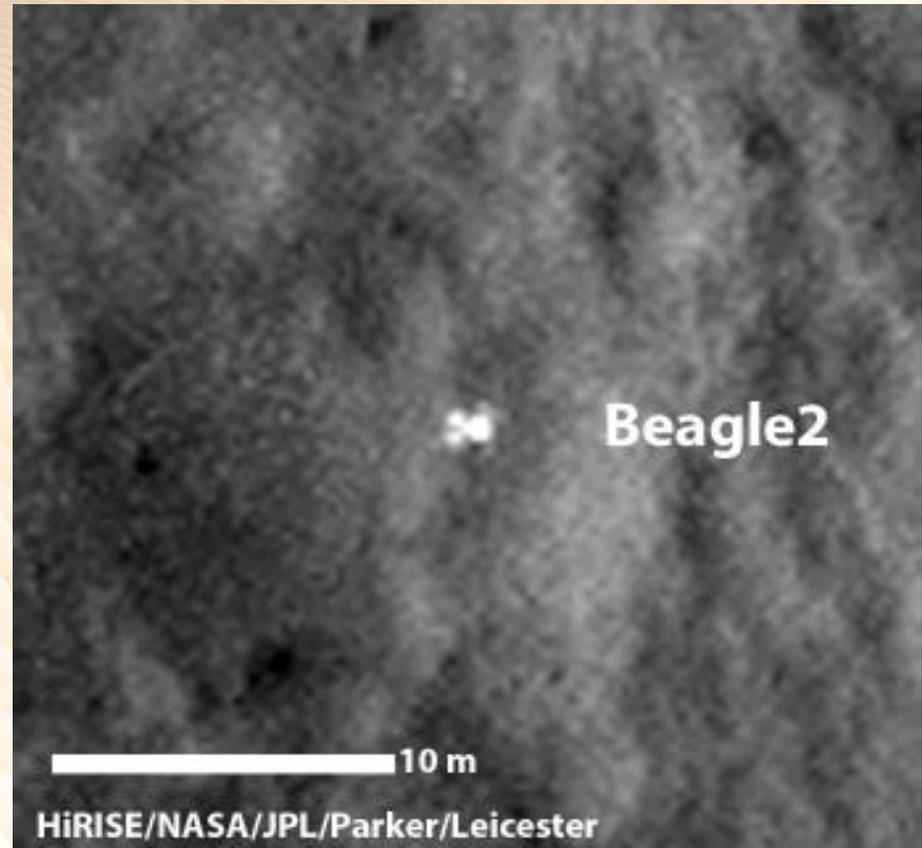


An injectite megapipe (conical hill). Image is ~1 km across. HiRISE image PSP\_001641\_1735

# Beagle-2 Lander Found on Mars

[http://www.esa.int/Our\\_Activities/Space\\_Science/Mars\\_Express/Beagle-2\\_lander\\_found\\_on\\_Mars](http://www.esa.int/Our_Activities/Space_Science/Mars_Express/Beagle-2_lander_found_on_Mars)

- *The UK-led Beagle-2 Mars lander, which was lost on Mars in late 2003, has been found in images taken by MRO HiRISE*
- *The lander is seen partially deployed on the surface, showing that the entry, descent and landing sequence worked and it did indeed land on Mars on Christmas Day 2003*
- *The high resolution images were initially searched by Michael Croon, a former member of the Mars Express operations team at ESA's Space Operations Centre, ESOC, working in parallel with members of the Beagle-2 industrial and scientific teams.*



This close-up image shows the lander in a partially deployed configuration, with at most three of the four solar panels open. Beagle-2, lies within its expected landing area at a distance of about 5 km from its center.

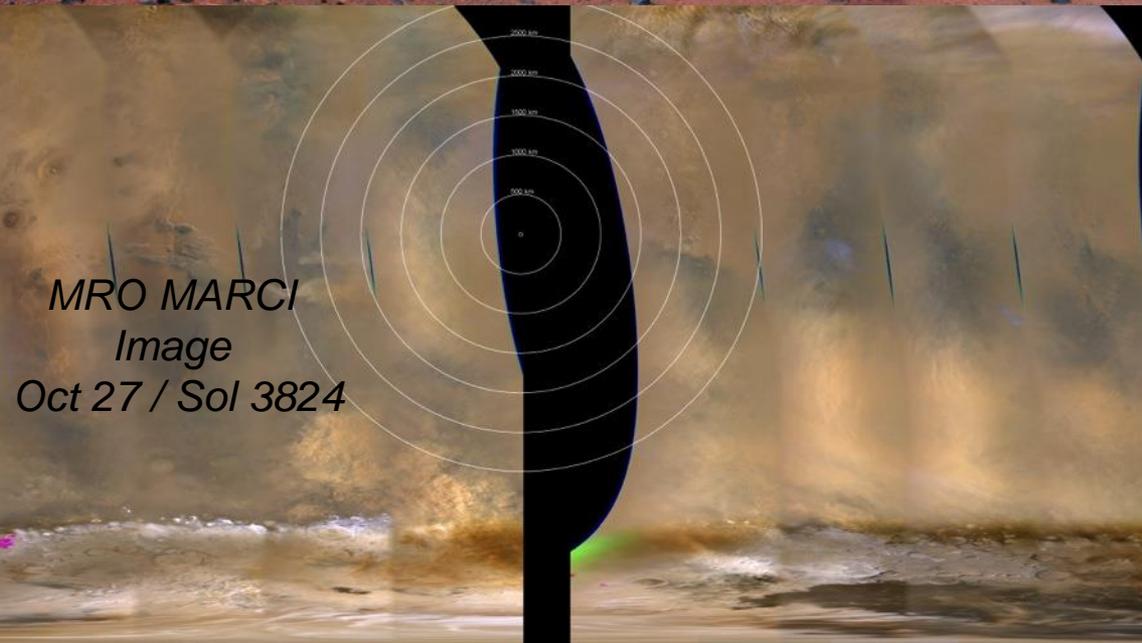


# Wdowiak Ridge

Pancam acquired on Sol 3786 showing Wdowiak Ridge and Endeavour Crater

## Regional Dust Storm

Dust opacity ( $\tau$ ) spiked to  $>2$  on Sol 3823, back down to 1.6 during “A-storm”, first of 3 typical dust events during southern spring and summer



MRO MARCI  
Image  
Oct 27 / Sol 3824

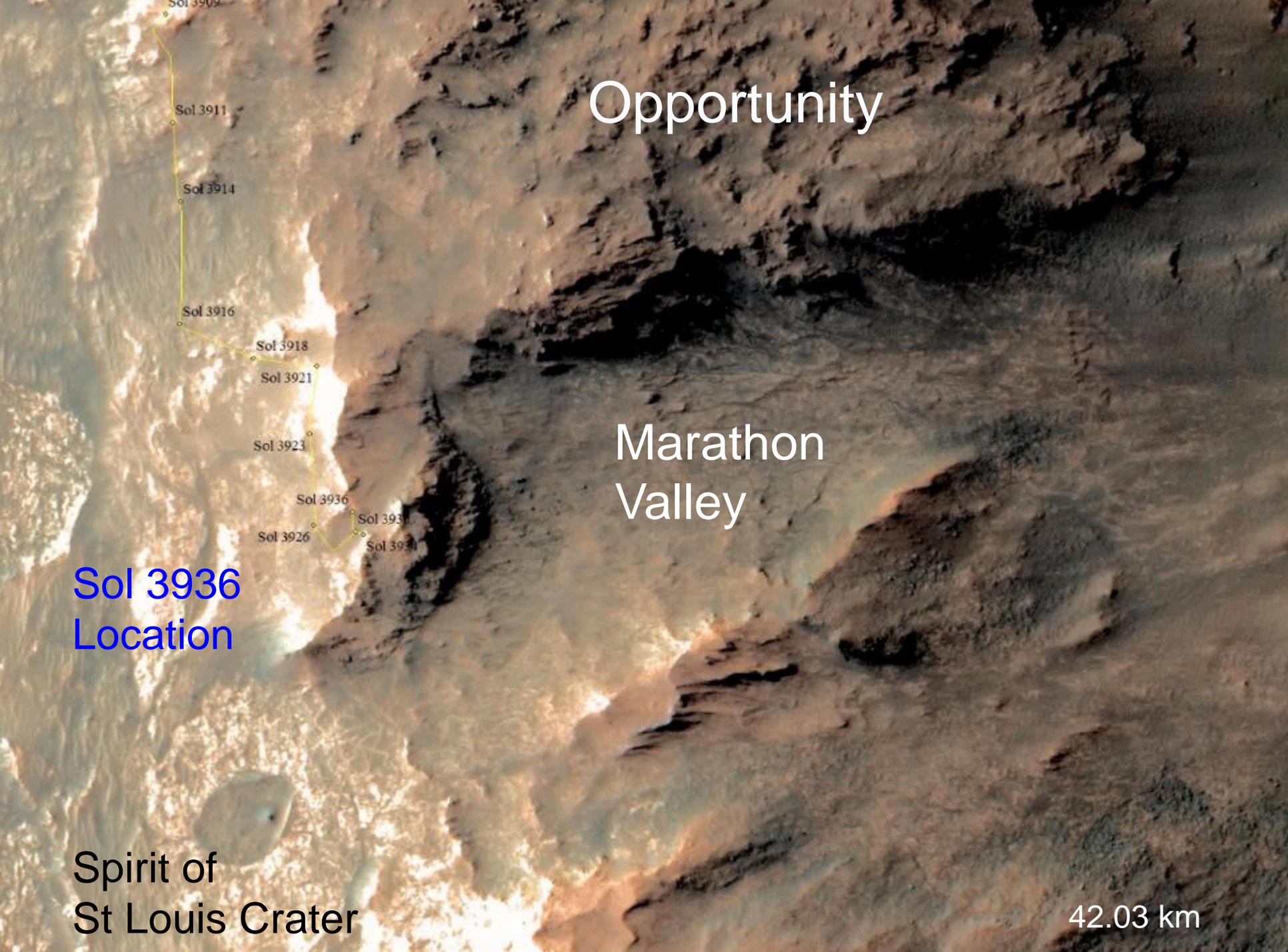
Opportunity

Marathon  
Valley

Sol 3936  
Location

Spirit of  
St Louis Crater

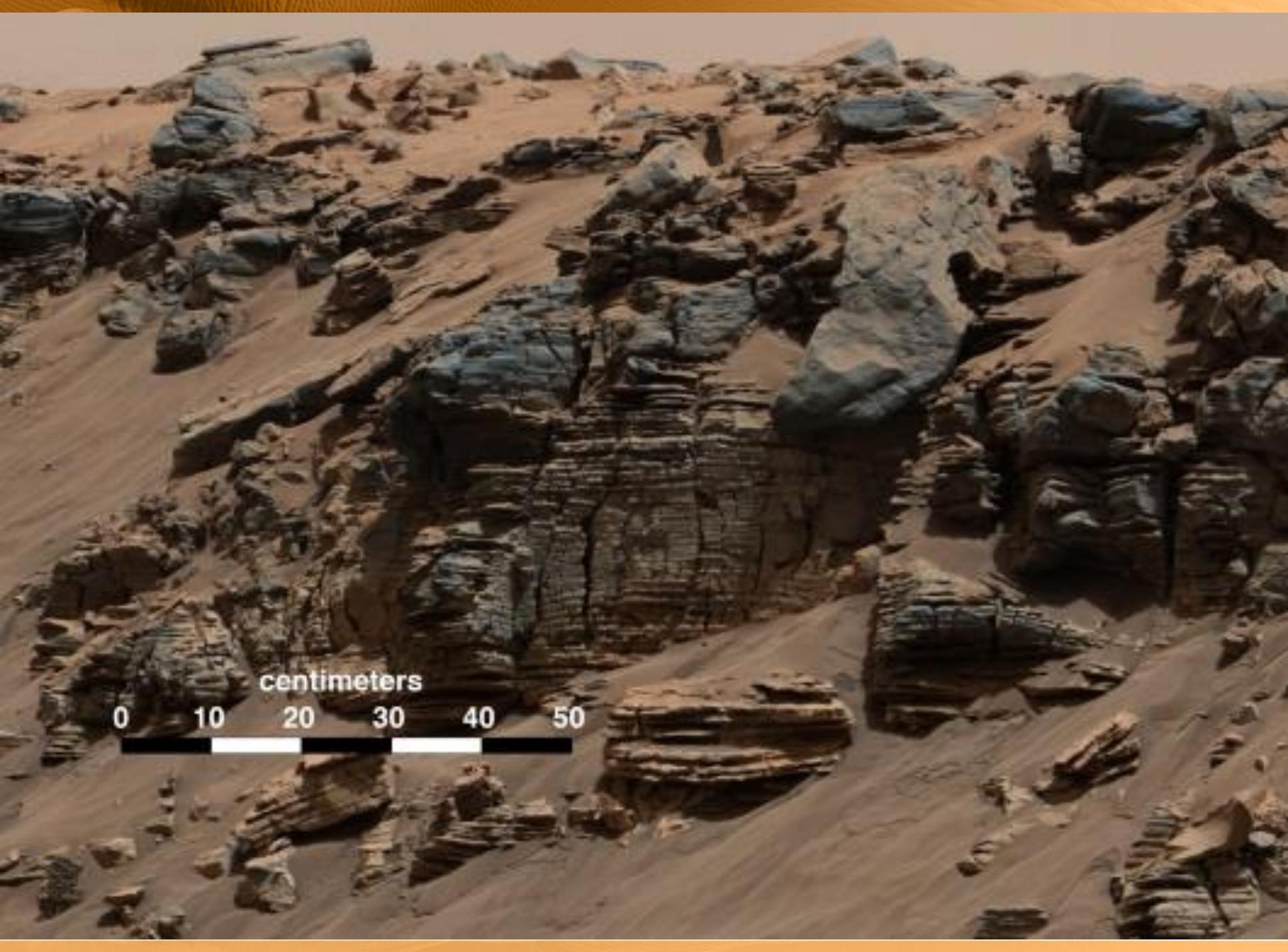
42.03 km

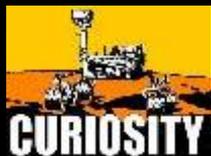
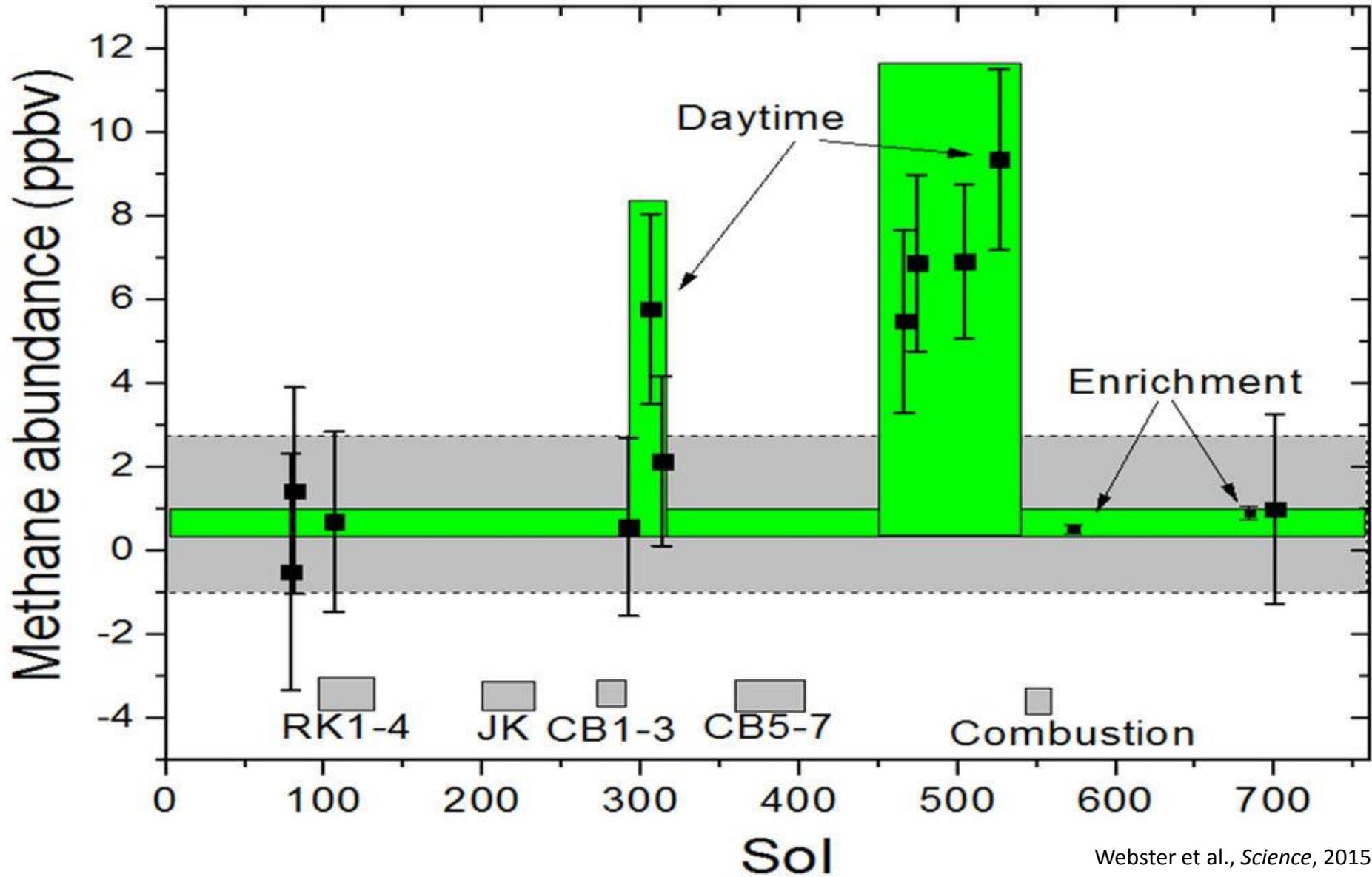


# Summit Cape Tribulation

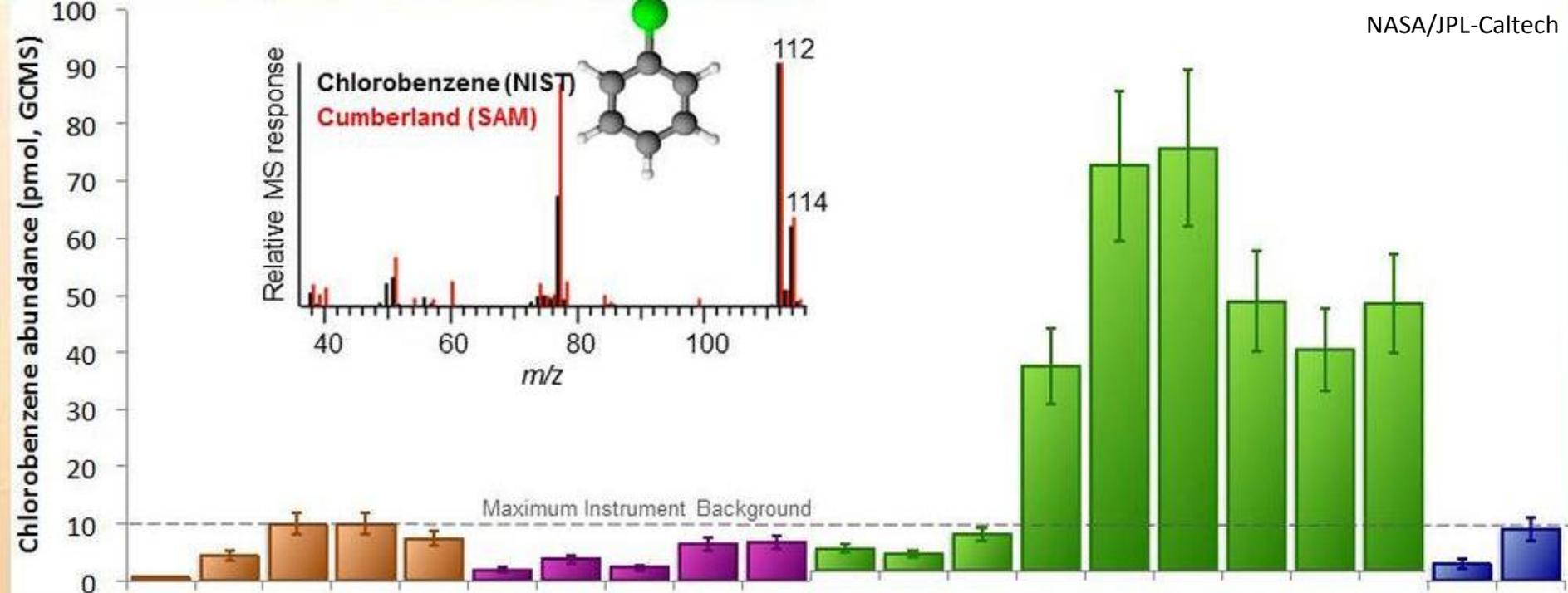


Pancam Sol 3902 Looking South to Marathon Valley





Curiosity measured a background methane abundance of 0.7 ppbv and a ten-fold enhancement that lasted ~ 60 sols



ROCKNEST

JOHN KLEIN

CUMBERLAND

CONFIDENCE  
HILLS

**Chlorobenzene was detected above the instrument background in the Cumberland drilled sample. The chlorine likely is derived from perchlorate in the sedimentary rock.**



**Aeolis Mons  
(Mount  
Sharp)**

Salsberry Peak

Whale Rock

Gilbert Peak

Chinle

Alexander Hills

Book Cliffs

Comb Ridge

Pink Cliffs

Confidence Hills

Mid-Drive  
Observation  
**Longer Stop**

NASA/JPL-Caltech/MSSS

Curiosity is completing a 5-month campaign at Pahrump Hills, the first outcrop of bedrock that composes the base of Mount Sharp. A first pass through the 10-m section used imagers and ChemCam to survey the morphology and chemistry.

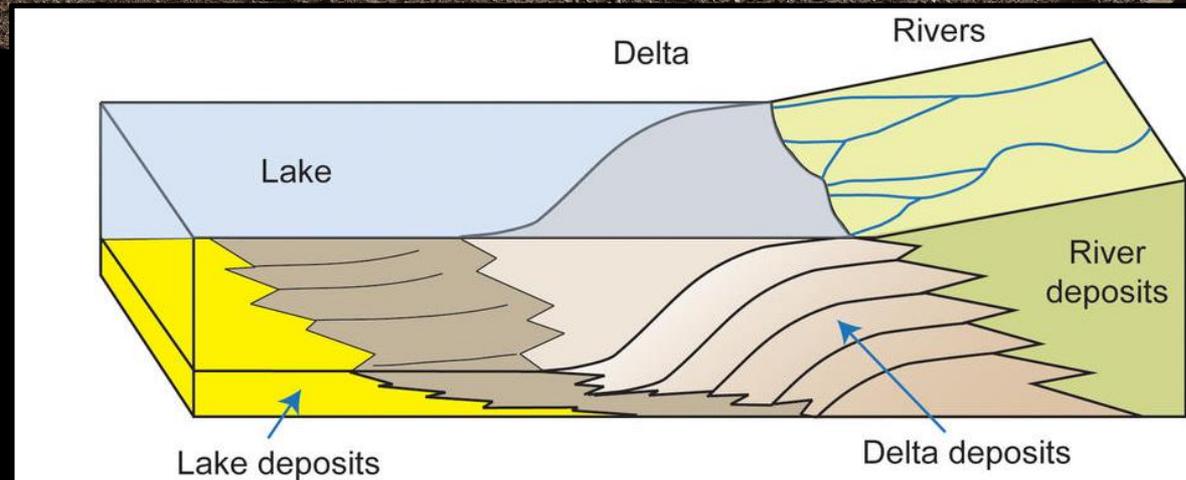


← Mount Sharp

meters  
0 2 4 6 8 10

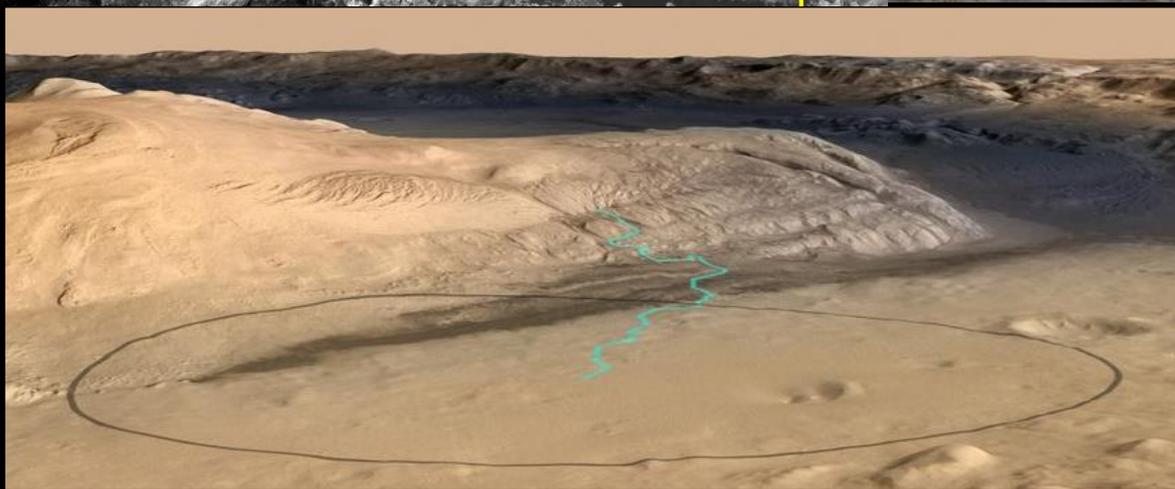
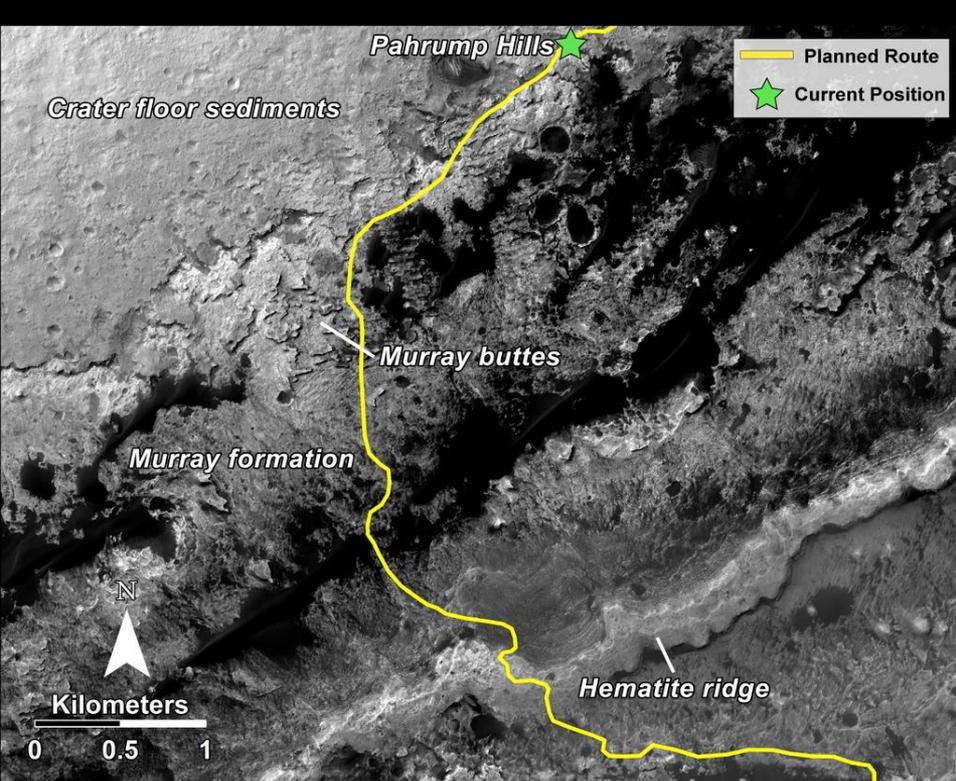
centimeters  
0 20 40 60 80 100

Possible Lake Deposits?



Curiosity is exploring a hypothesis that southward-tilted sandstone beds on Gale Crater's plains indicate fluvial transport of sediment toward Mount Sharp, building up lake deposits there.





NASA/JPL-Caltech/MSSS

Curiosity's Extended Mission will explore Mt. Sharp, with an emphasis on understanding the subset of habitable environments that preserve organic carbon



#JOURNEYTOMARS